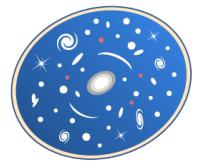


# The Reionization Lensing Clusters Survey: Strong Lensing Analysis



Rachel Paterno-Mahler  
University of Michigan  
Clusters 2017  
July 4, 2017



**RELICS**  
Reionization Lensing Cluster Survey



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Benedetta Vulcani  
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University of Arizona 3/3/16

# RELICS: Reionization Lensing Cluster Survey

## Observations

190 orbits + 77 parallel (incl. 20 for SN follow-up)  
46 fields lensed by 41 clusters  
- 3 orbits ACS (minus archival)  
- 2 orbits WFC3/IR      Frontier Fields filters

## Science

high-redshift galaxies  
cluster mass scaling relations  
merger physics + DM constraints  
supernovae

## Delivery

no proprietary period HST images  
reduced images + catalogs  
2-3 months after completion of each field  
final high-z candidates + lens models  
Nov. 2017 (JWST GO call for proposals)



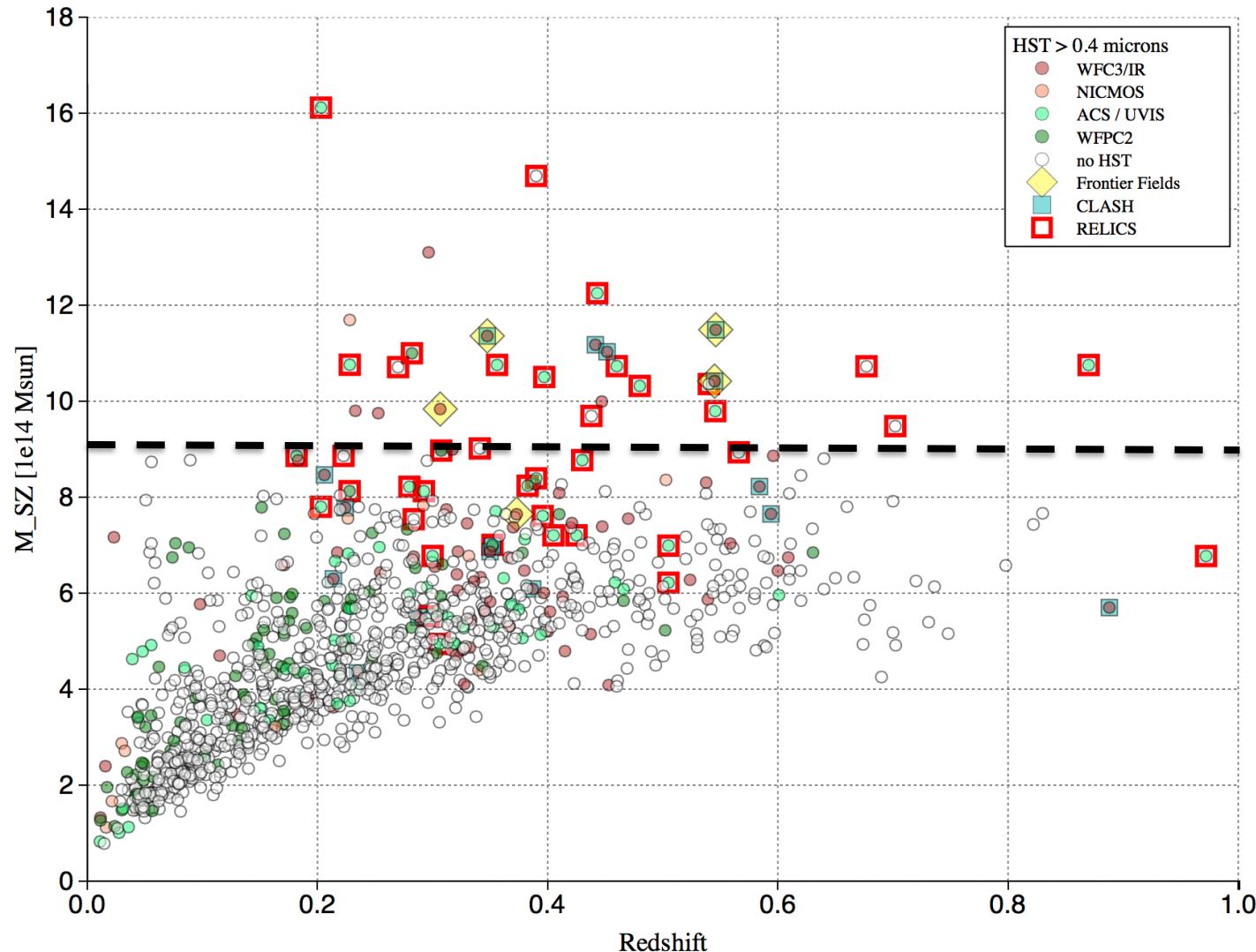
Abell 2537

(Slide courtesy D. Coe)

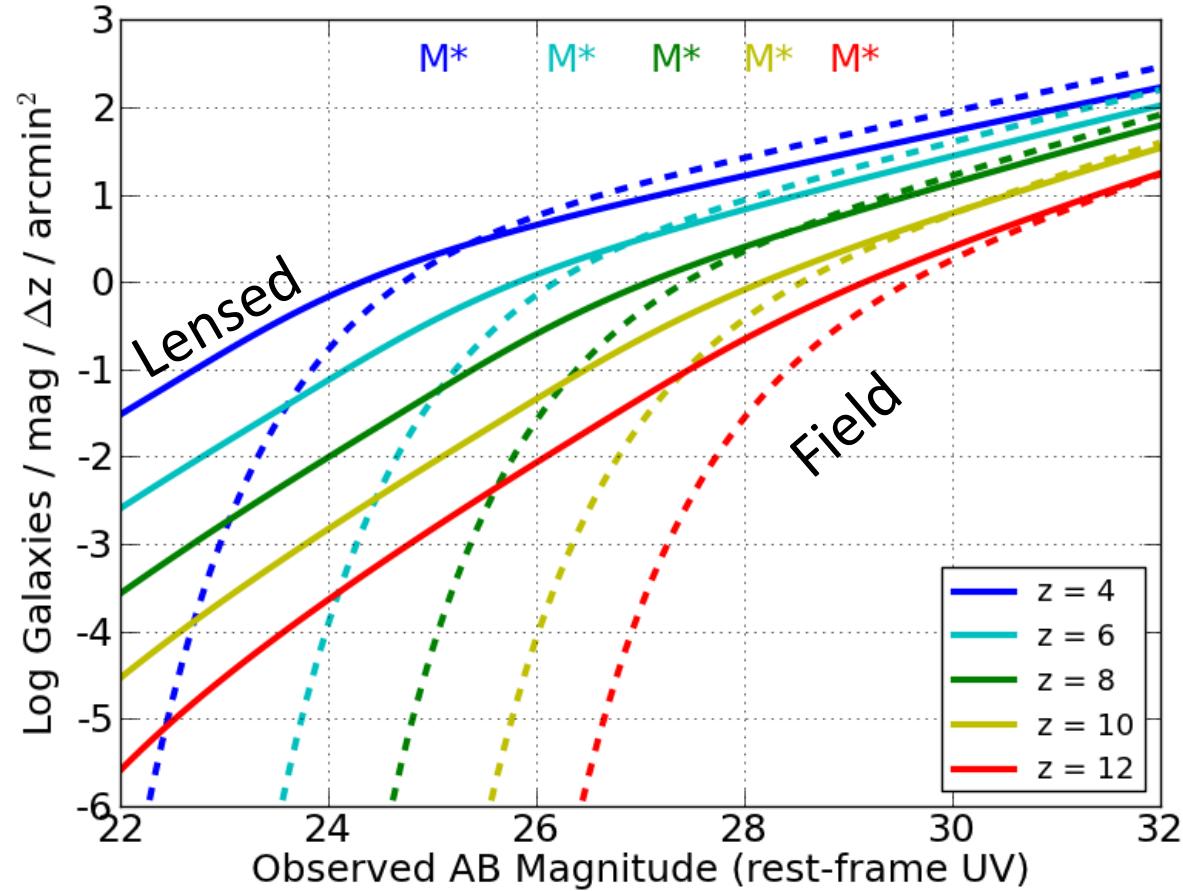
# Observations

- Hubble: 190 Orbits (PI: Coe)
- Spitzer: 100 Hours (PI: Bradac); 290 Hours DDT (PI: Soifer)
- Magellan Megacam: 1 Night (PI: Jones)
- Archival X-ray
- Contributions from existing programs:
  - Magellan LDSS3 (Sharon)
  - Keck DEIMOS (Bradac)
  - Subaru HSC (Dawson, MC3PO)

# Cluster Selection



# Cluster Lensing Enables Efficient Discovery of Distant Galaxies



Bouwens12a luminosity function  
evolving all 3 parameters

(Slide Courtesy D. Coe)

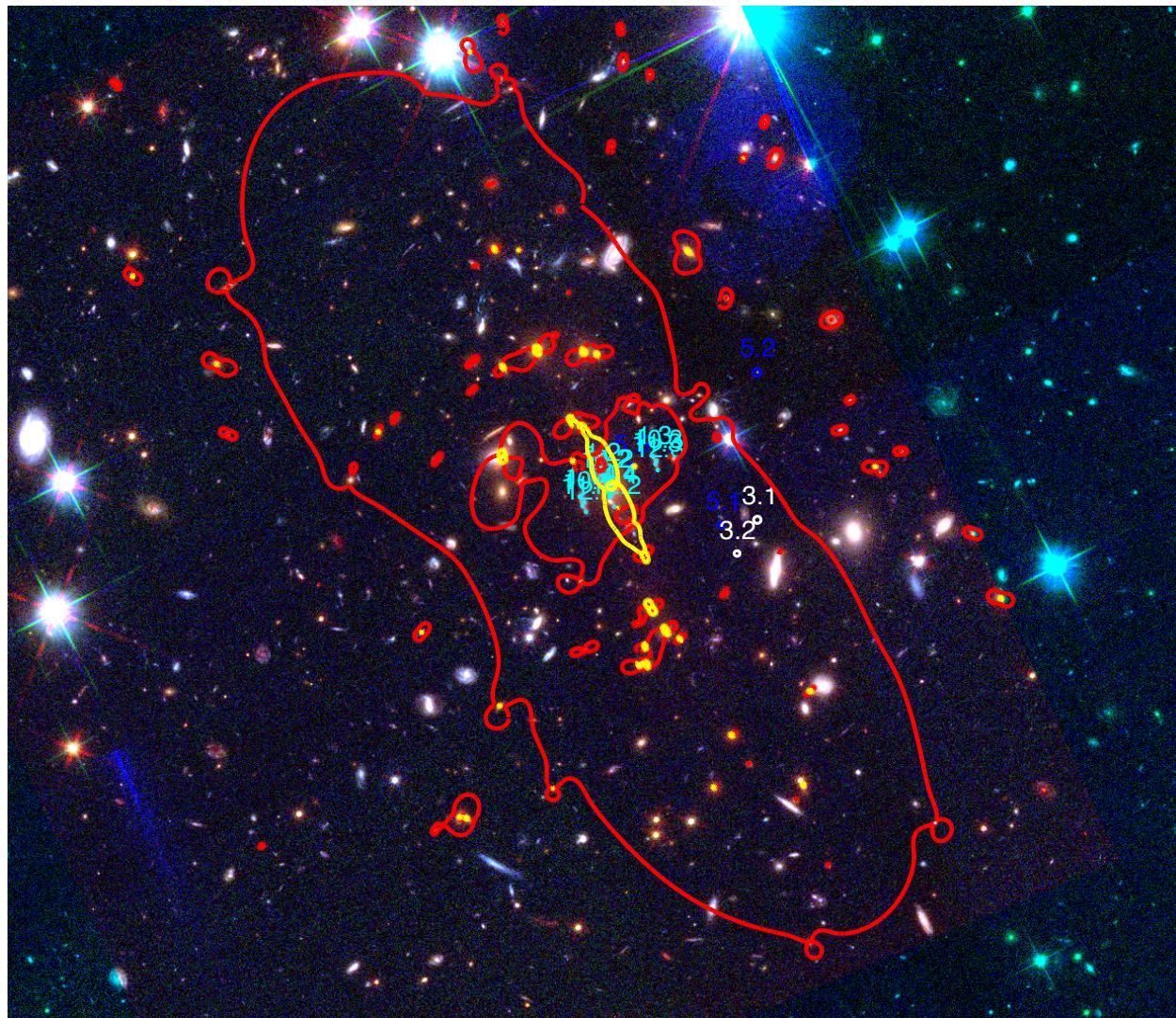
# Lens Modeling

- Whenever possible, at least one spectroscopic redshift (see Johnson & Sharon 2016)
- Models computed with Lenstool, which is a parametric lens modeling code
- Each halo modeled as a PIEMD
- Proceed iteratively until the model converges

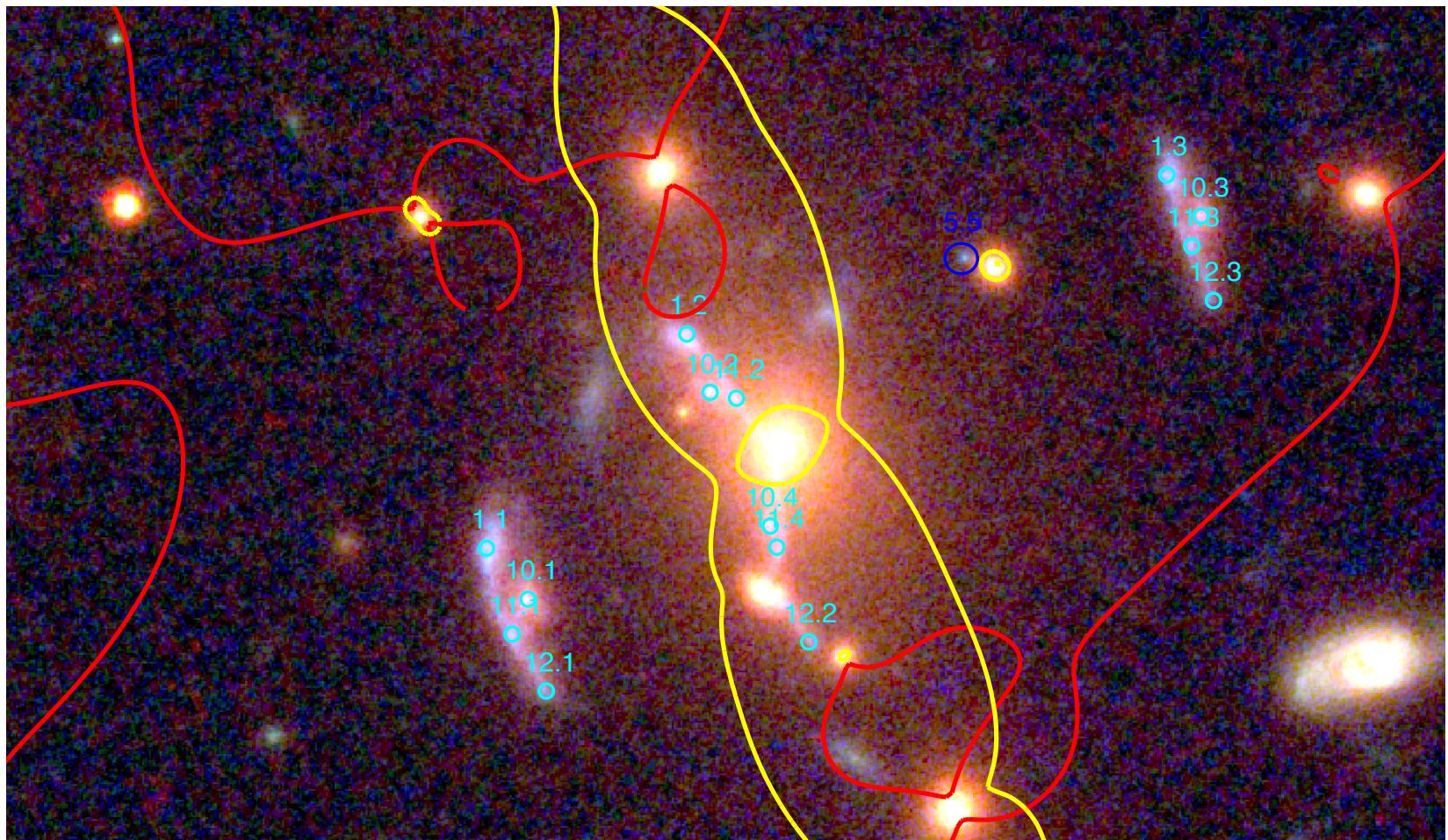
# Arc Selection



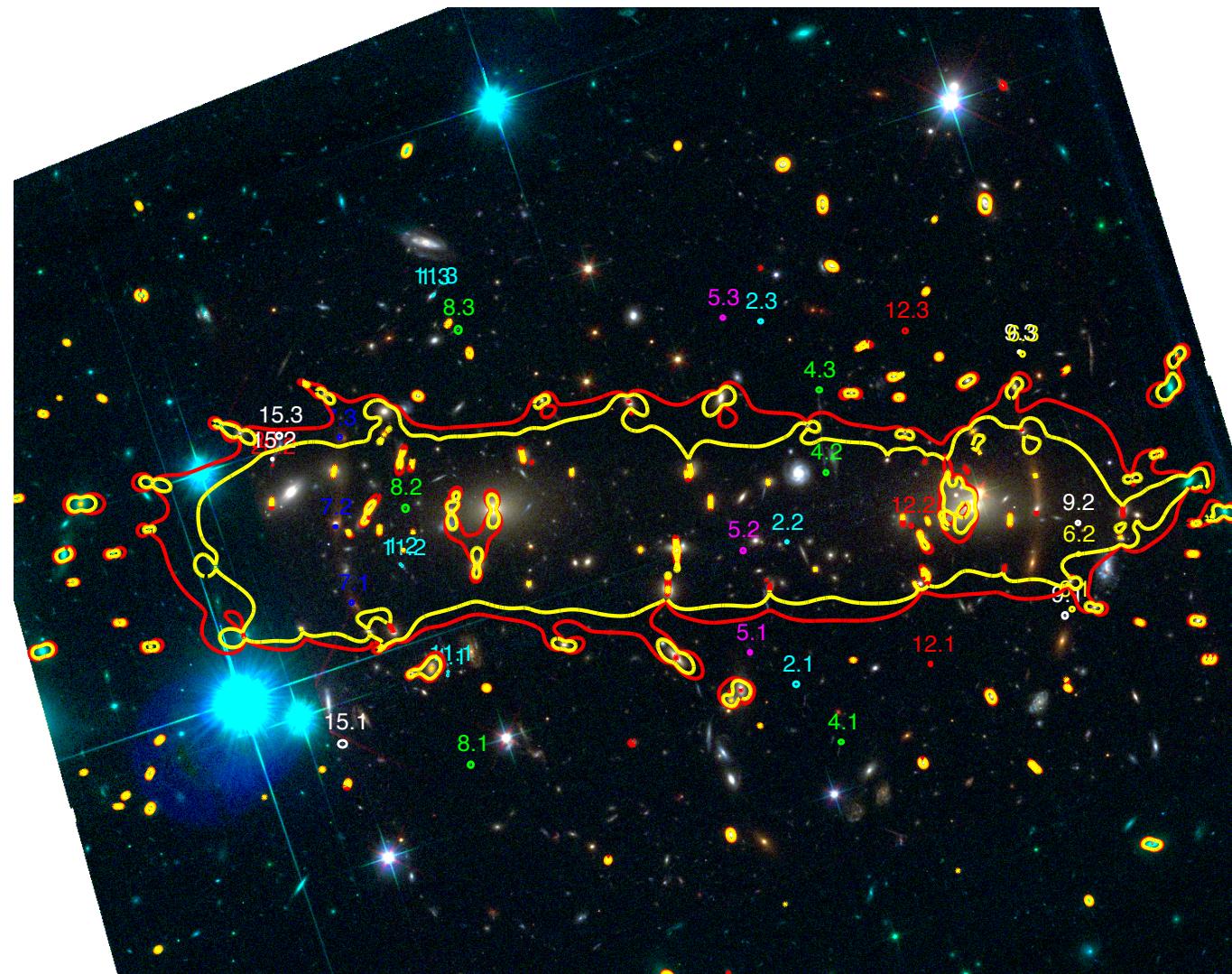
# SPT0615-57



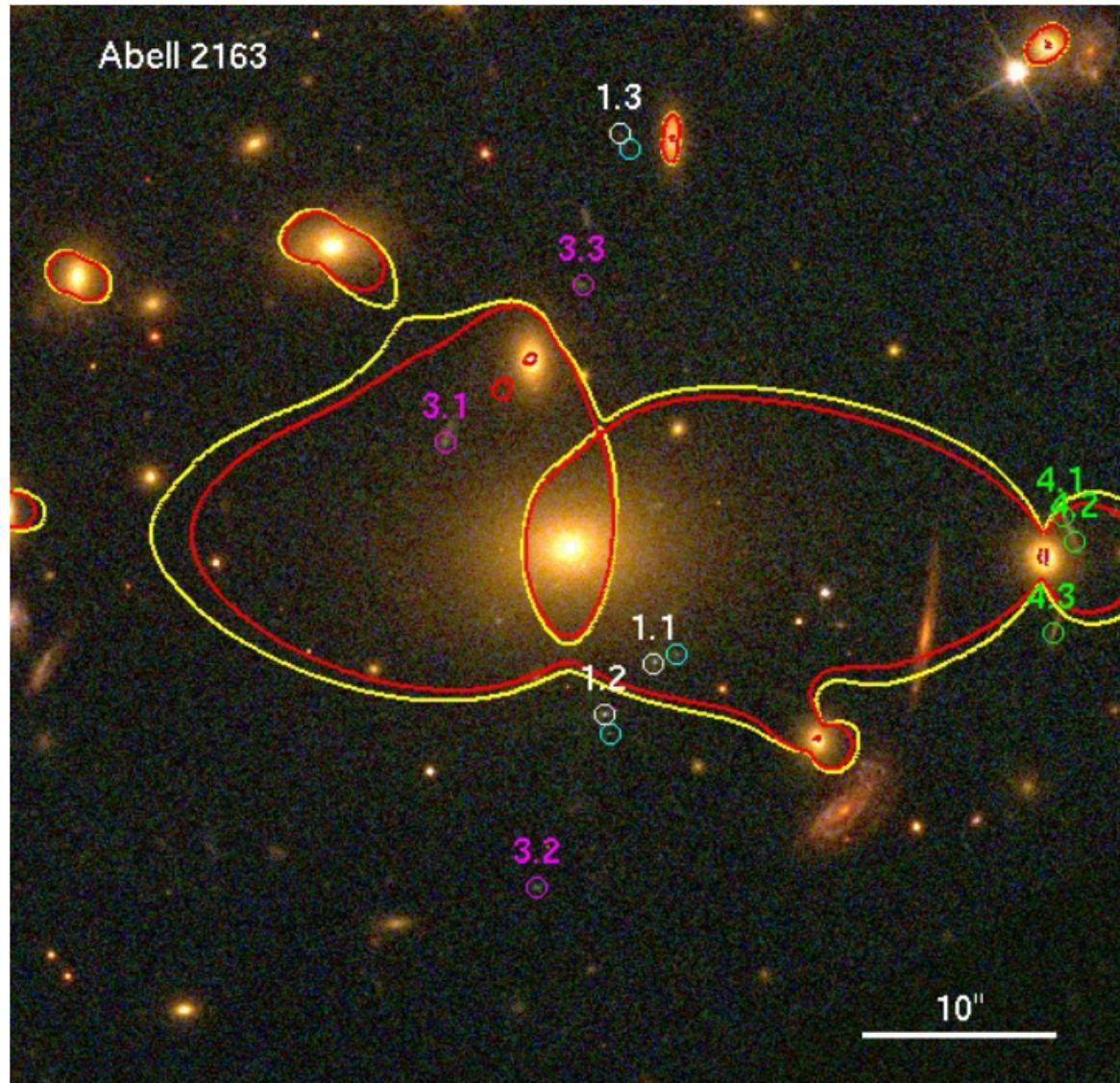
# SPT0615 Zoom



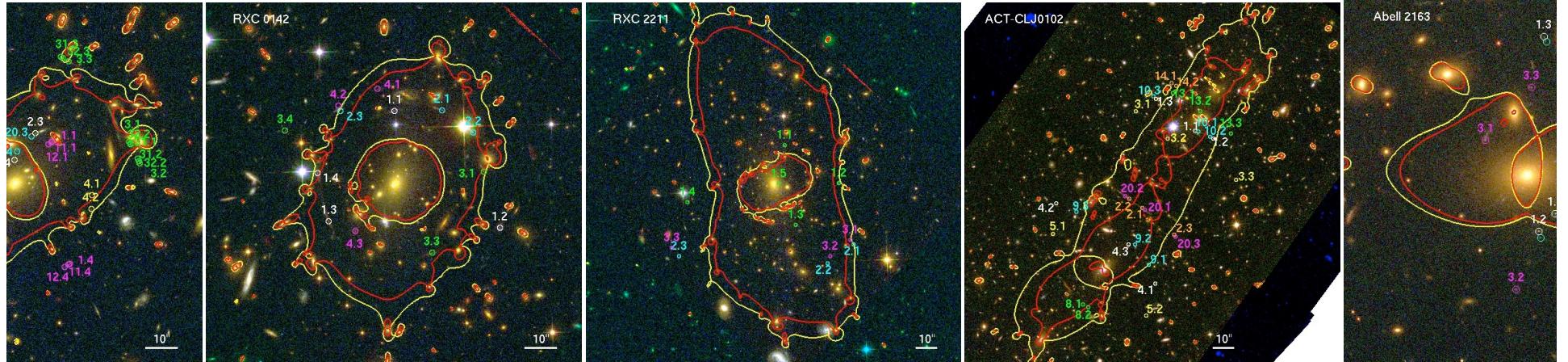
# MACS0556



# Abell 2163



Cerny et al., In Prep



## RELICS: STRONG LENS MODELS FOR FIVE GALAXY CLUSTERS FROM THE REIONIZATION LENSING CLUSTER SURVEY

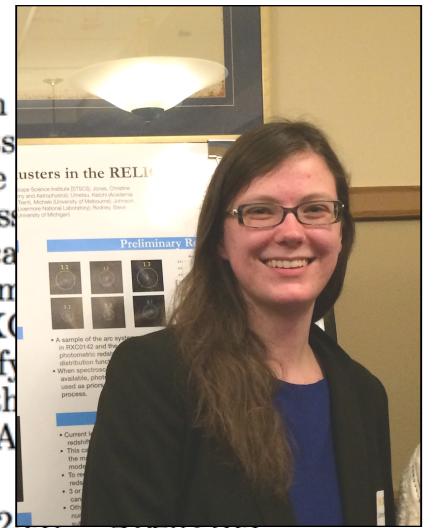
CATHERINE CERNY<sup>1</sup>, KEREN SHARON<sup>1</sup>, DAN COE<sup>2</sup>, ROBERTO AVILA<sup>2</sup> MARUSA BRADAC<sup>7</sup>, LARRY BRADLEY<sup>2</sup>, NICOLE CZAKON<sup>4</sup>, WILLIAM DAWSON<sup>8</sup>, TRACI JOHNSON<sup>1</sup>, CHRISTINE JONES<sup>3</sup>, RAMESH MAINALI<sup>5</sup> RACHEL PATERNO-MAHLER<sup>1</sup>, BRETT SALMON<sup>2</sup> SARA OGAZ<sup>2</sup>, STEVE RODNEY<sup>9</sup>, DANIEL STARK<sup>5</sup>, LOU STROLGER<sup>2</sup>, MICHELE TRENTI<sup>6</sup>, KEIICHI UMETSU<sup>4</sup>, RELICS TEAM

*Draft version June 15, 2017*

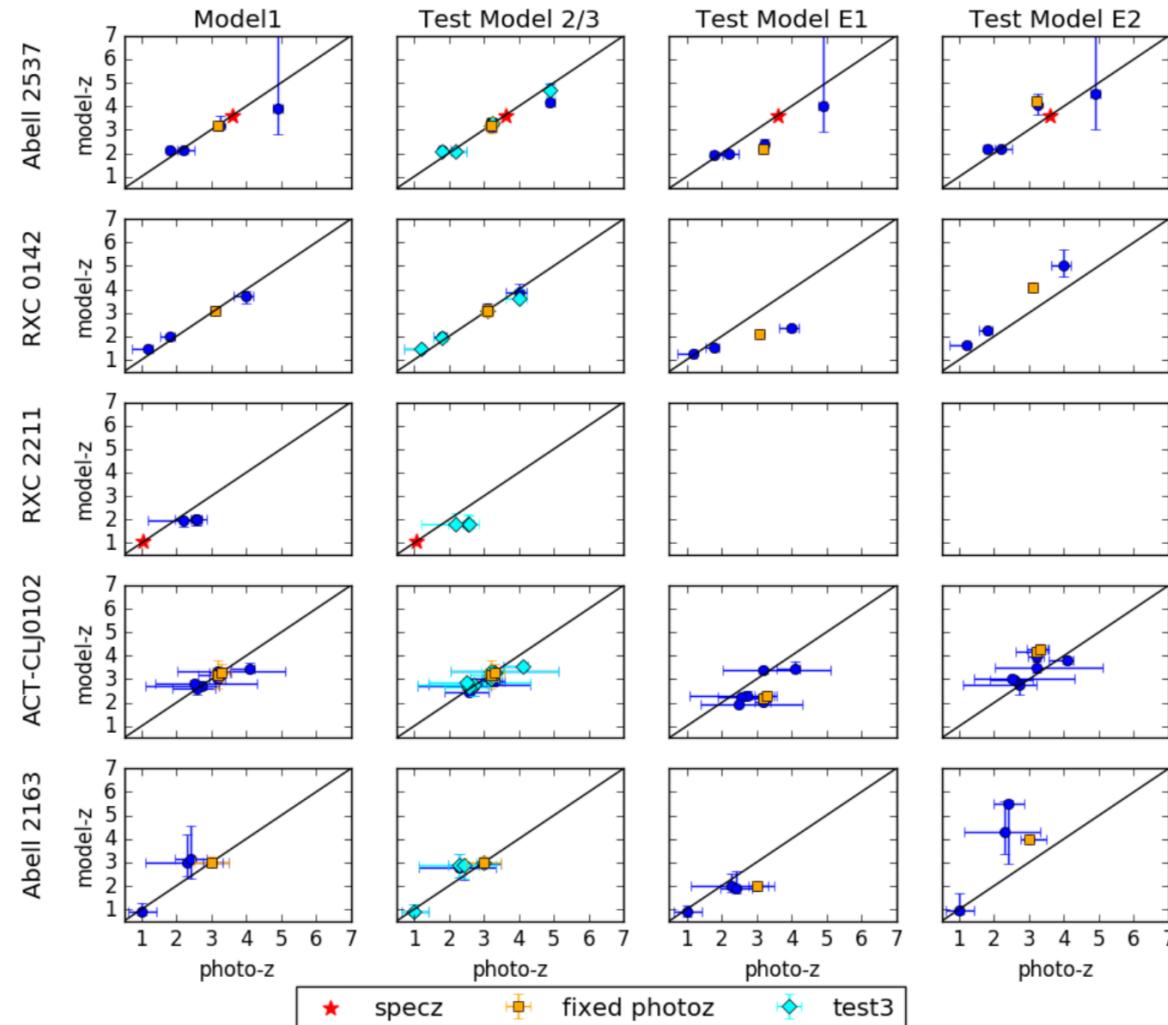
### ABSTRACT

We present strong lensing models for five galaxy clusters from the Reionization Survey (RELICS) Hubble Treasury Program. The strong gravitational lensing effects significantly magnify background galaxies, which enhances our ability to discover the high redshift galaxies at  $z \sim 9 - 12$  needed for a statistically significant sample necessary to constrain the galaxy luminosity function. The intrinsic properties of high-redshift galaxy can be calculated by accounting for the lensing effect. Here, we present the first five lens models for the RELICS program, based on new *HST* WFC3 and ACS imaging of the clusters RXC J0142.9+4438, RXC J2211.7–0349, Abell 2537, Abell 2163, RXC J2211.7–0349, and ACT–CLJ0102–49151. We quantify the uncertainty on the lensing magnification due to statistical and systematic errors related to the lens modeling process. The models are made available to the community through the Mikulski Archive for Space Telescopes.

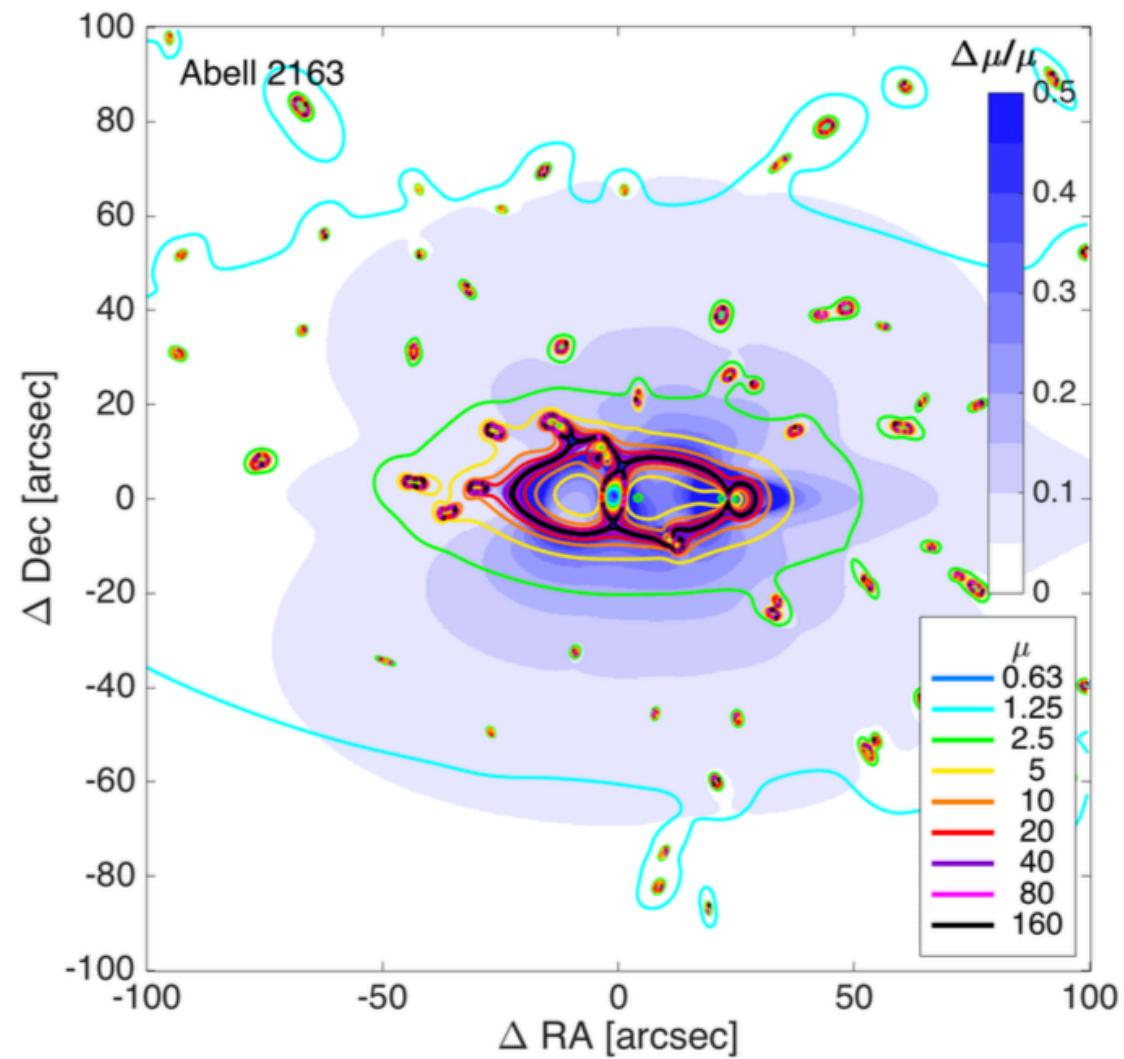
*Subject headings:* galaxies: clusters: individual (RXC J0142.9+4438, Abell2537, Abenz163, RXCJ2211.7-0349, ACT-CLJ0102-49151) — gravitational lensing: strong



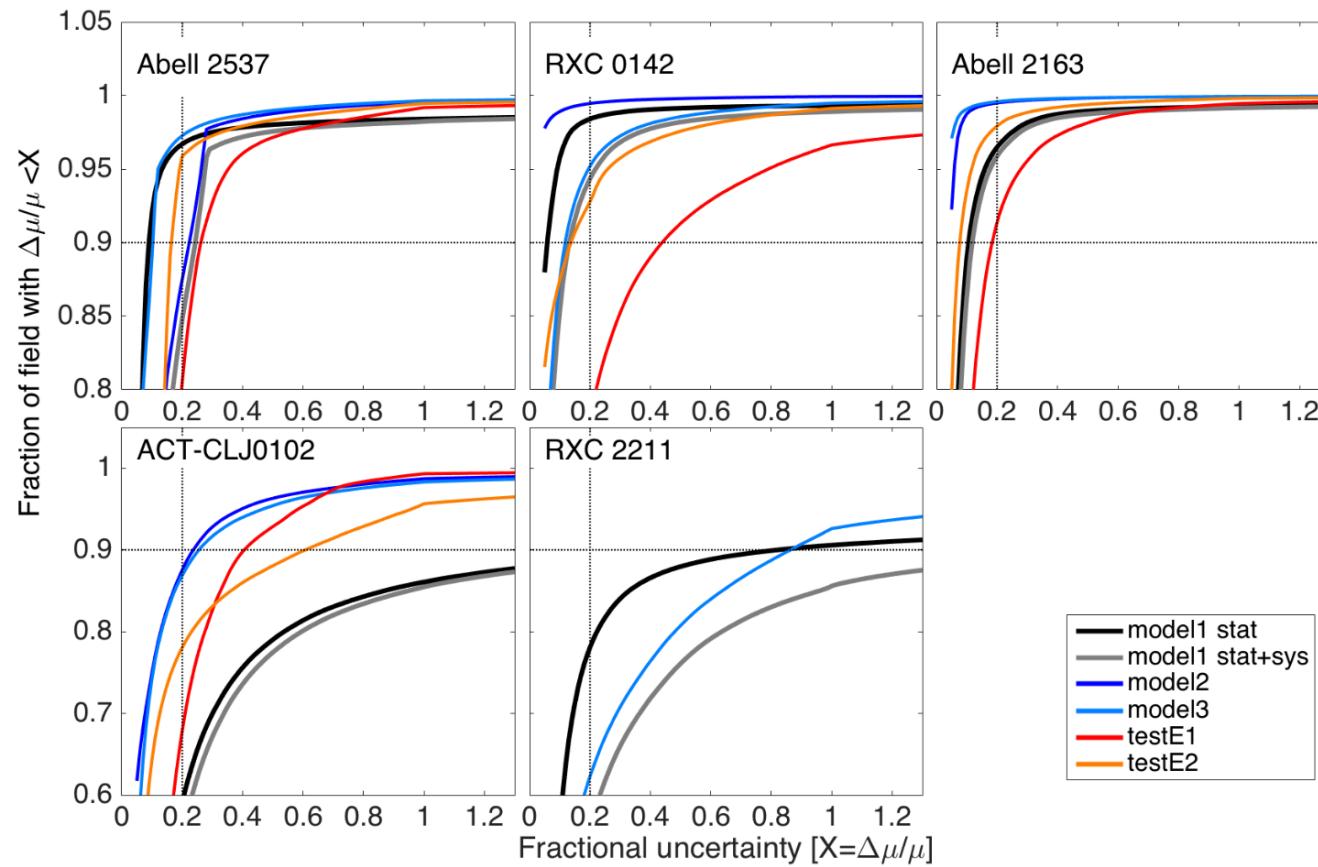
# What if there are no spectroscopic redshifts, I



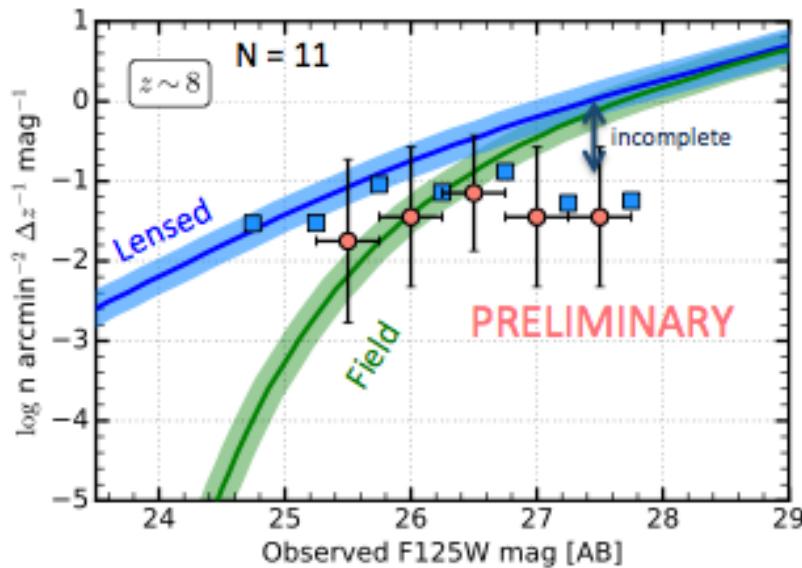
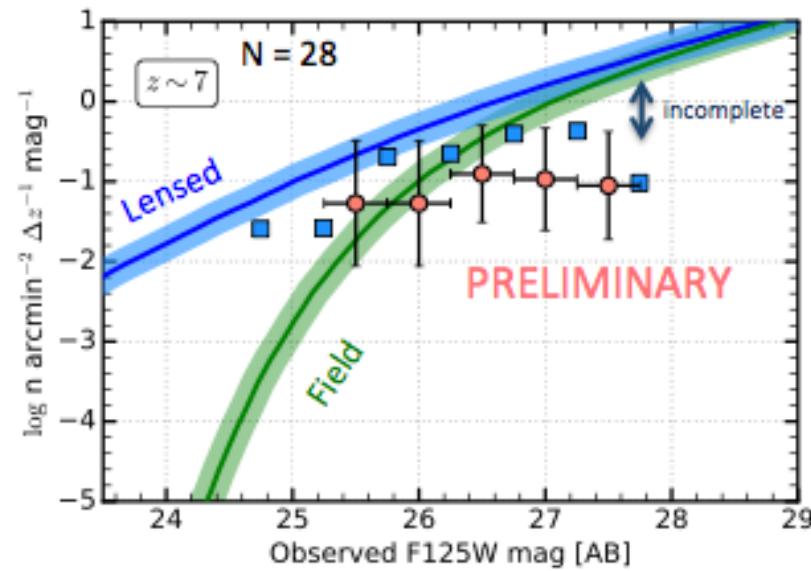
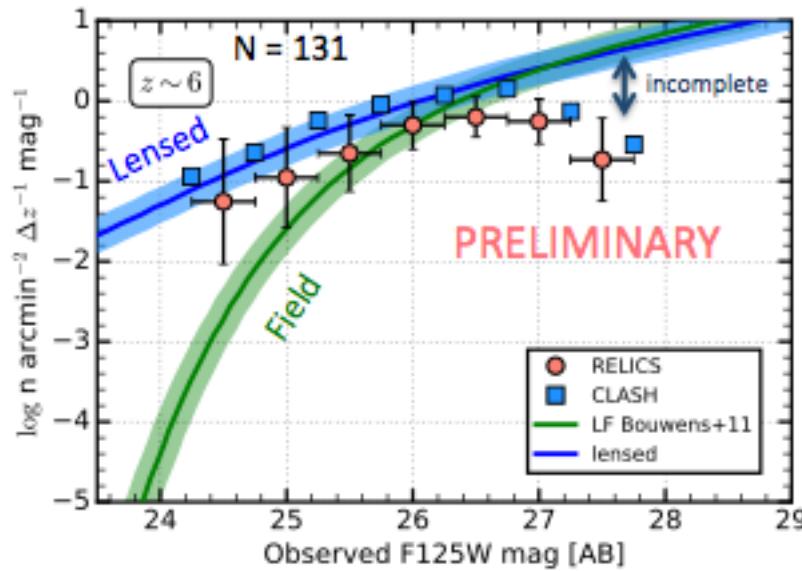
# Magnification Map



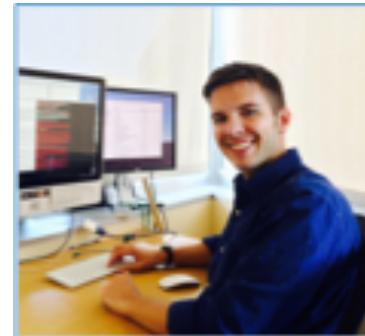
# What if there are no spectroscopic redshifts, II



# Preliminary High Redshift Candidates



Salmon et al. in prep.: 26 /  
41 RELICS clusters



# Future Work

- All models will be public in the fall
- Can do much cluster physics
  - Compare DM and X-ray centroid offsets
  - Explore DM – concentration relation
  - Scaling relations with the X-ray data