

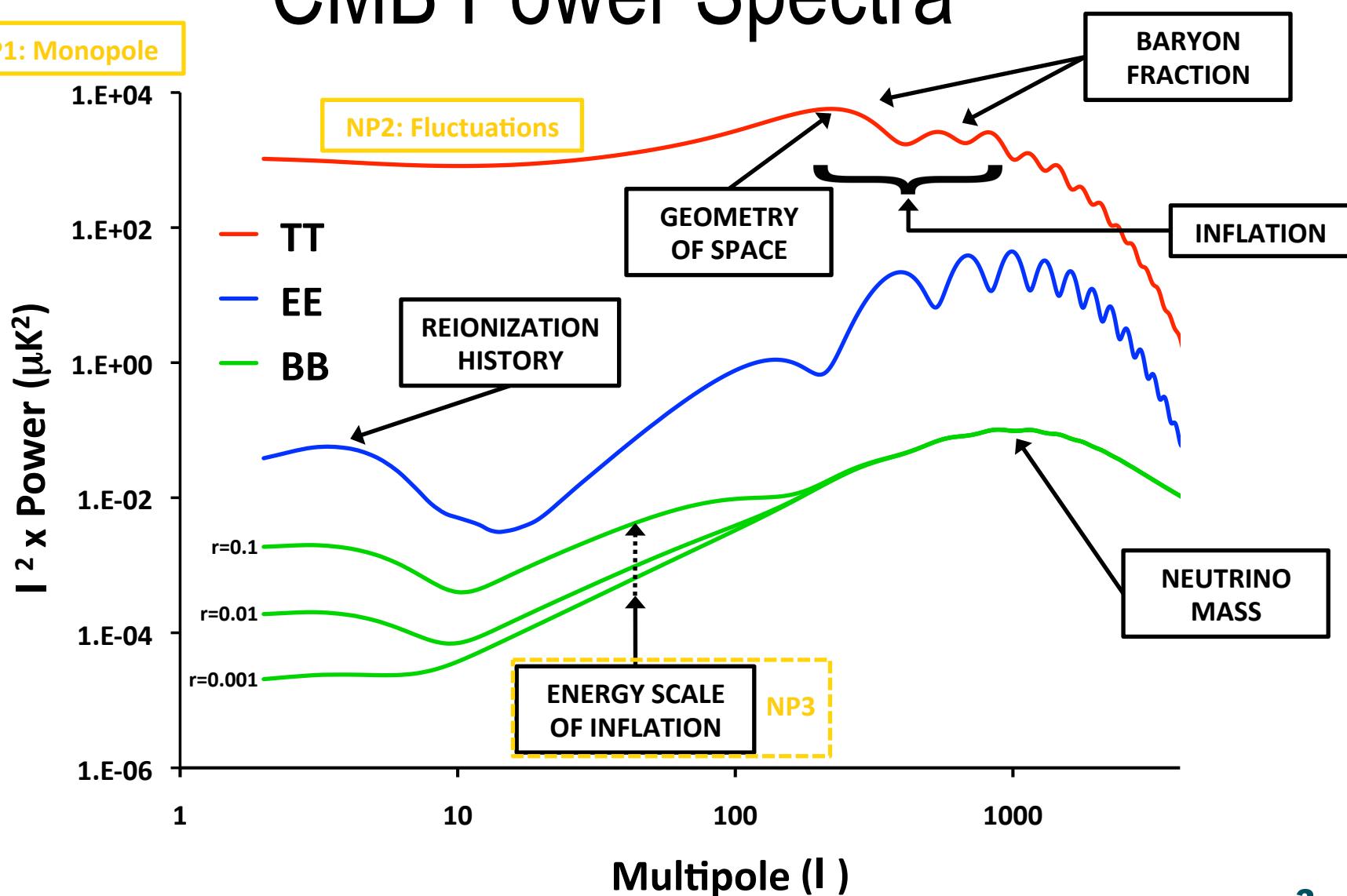
High Performance Computing For B-Mode Data Analysis

Julian Borrill

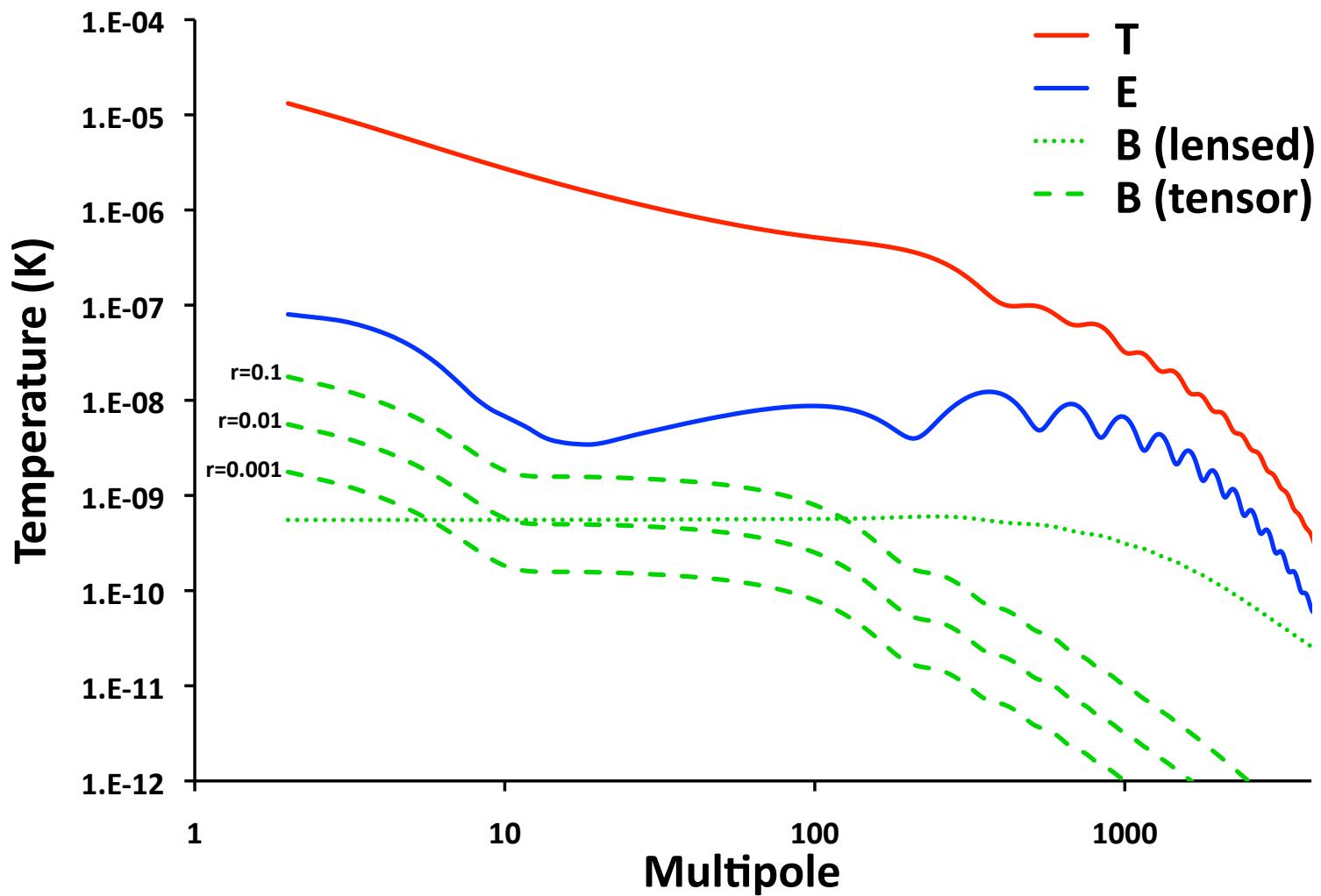
Computational Cosmology Center, Berkeley Lab
& Space Sciences Laboratory, UC Berkeley



CMB Power Spectra



Another View ...



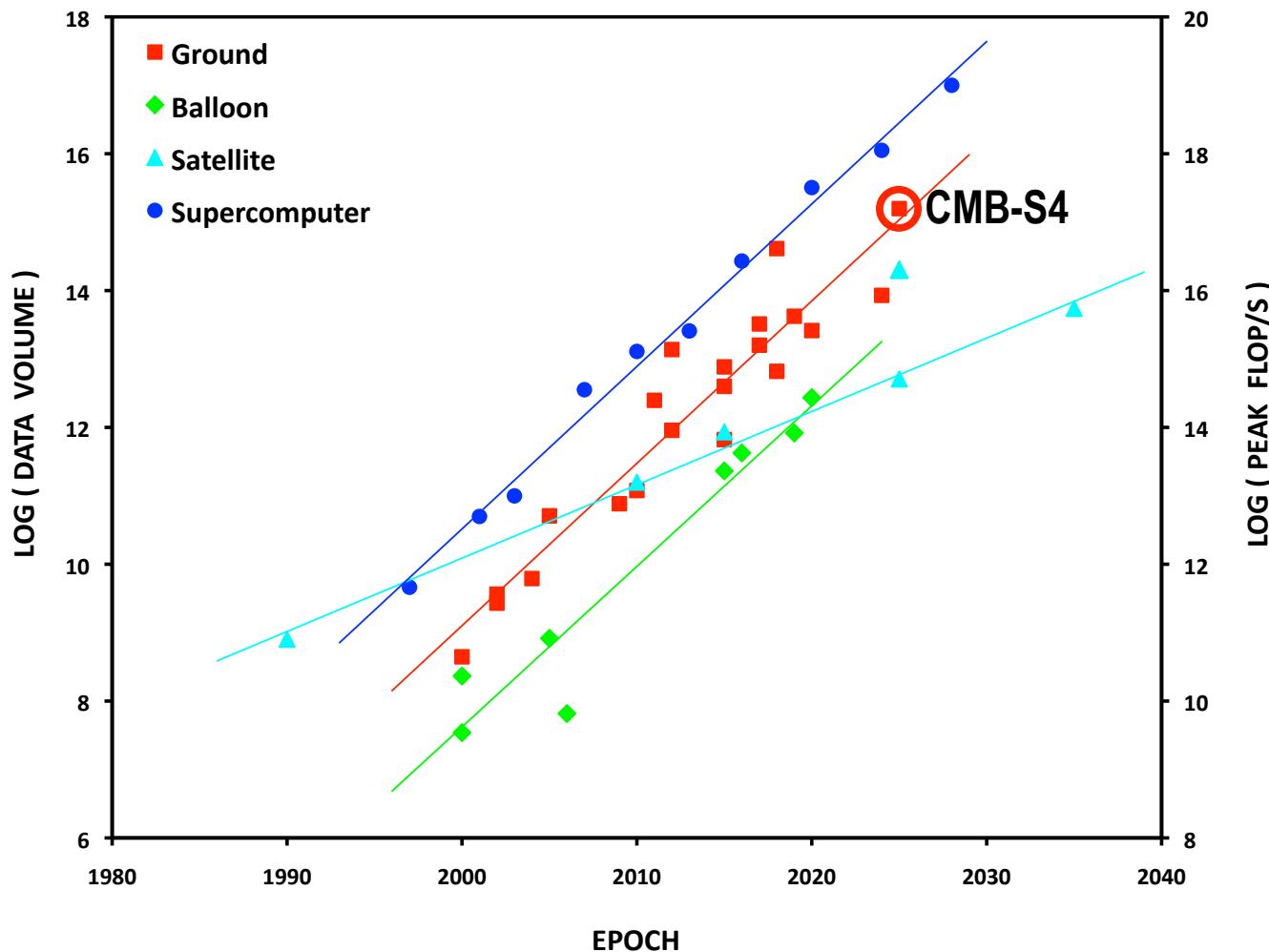
CMB Signals

| COMPONENT | AMPLITUDE (K) | ERA |
|---------------------------|---------------|---|
| TT : Monopole | 1 | 1968 (Penzias & Wilson) |
| TT : Anisotropy | 10^{-5} | 1990 (COBE) |
| TT : 1 st Peak | 10^{-6} | 2000 (BOOMERanG, MAXIMA) |
| EE : Reionization | 10^{-7} | 2005 (DASI) |
| BB : Lensing | 10^{-9} | 2015 (SPT, POLARBEAR) |
| BB : Gravity Waves | $< 10^{-9}$ | (AdvACTpol, BICEP3, CLASS, Simons Array, SPT-3G, CMB-S4; EBEX-10K, PIPER, SPIDER-2; LiteBIRD, PIXIE, COrE+) |

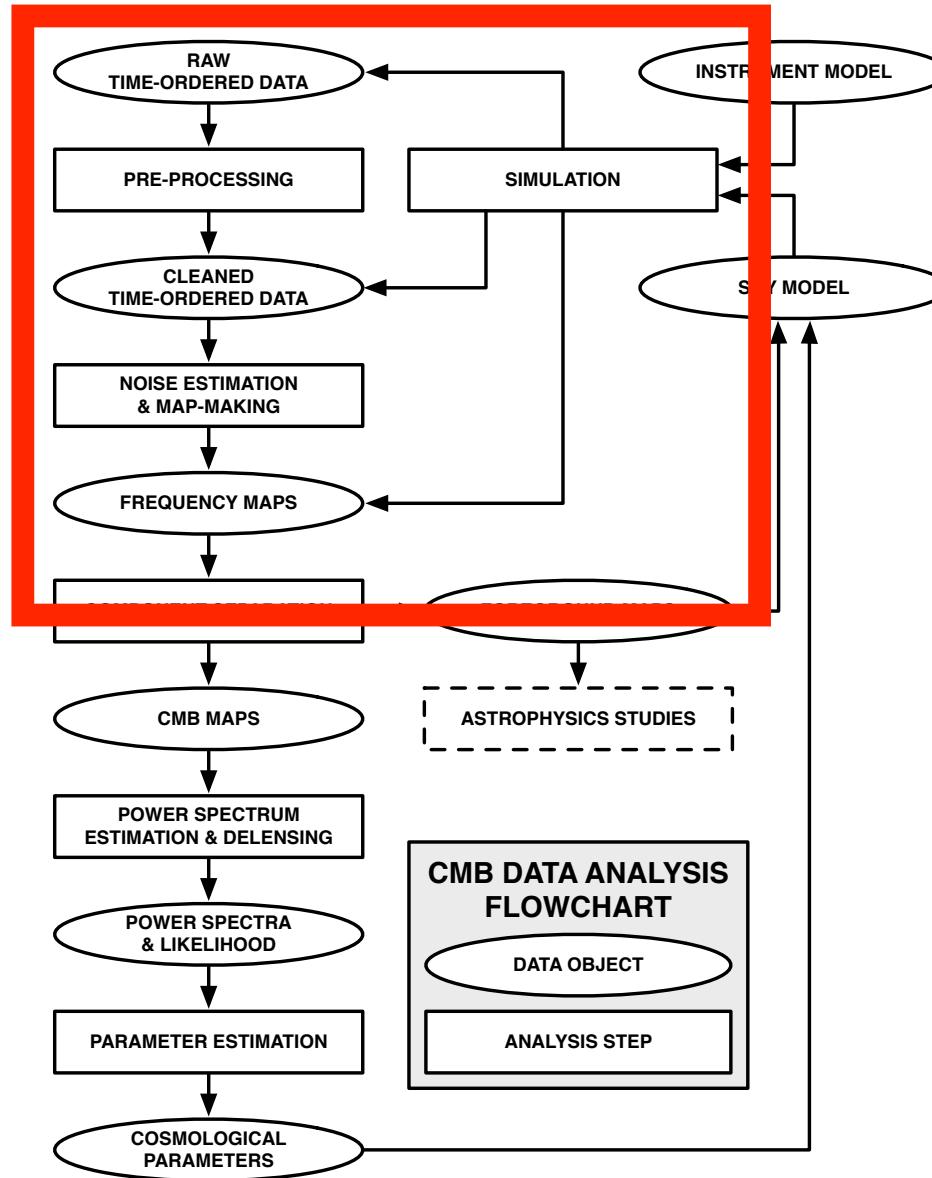
CMB Data Evolution

- Detecting fainter signals requires
 - Larger data sets to beat down noise
 - Tighter control of systematics
 - Environment
 - Foregrounds, Atmosphere
 - Instrument
 - Beams, Bandpasses, Noise, (HWP)
 - Analysis
 - Approximations, Biases, Uncertainties

CMB Data Growth



CMB Data Flow



CMB Data Algorithms

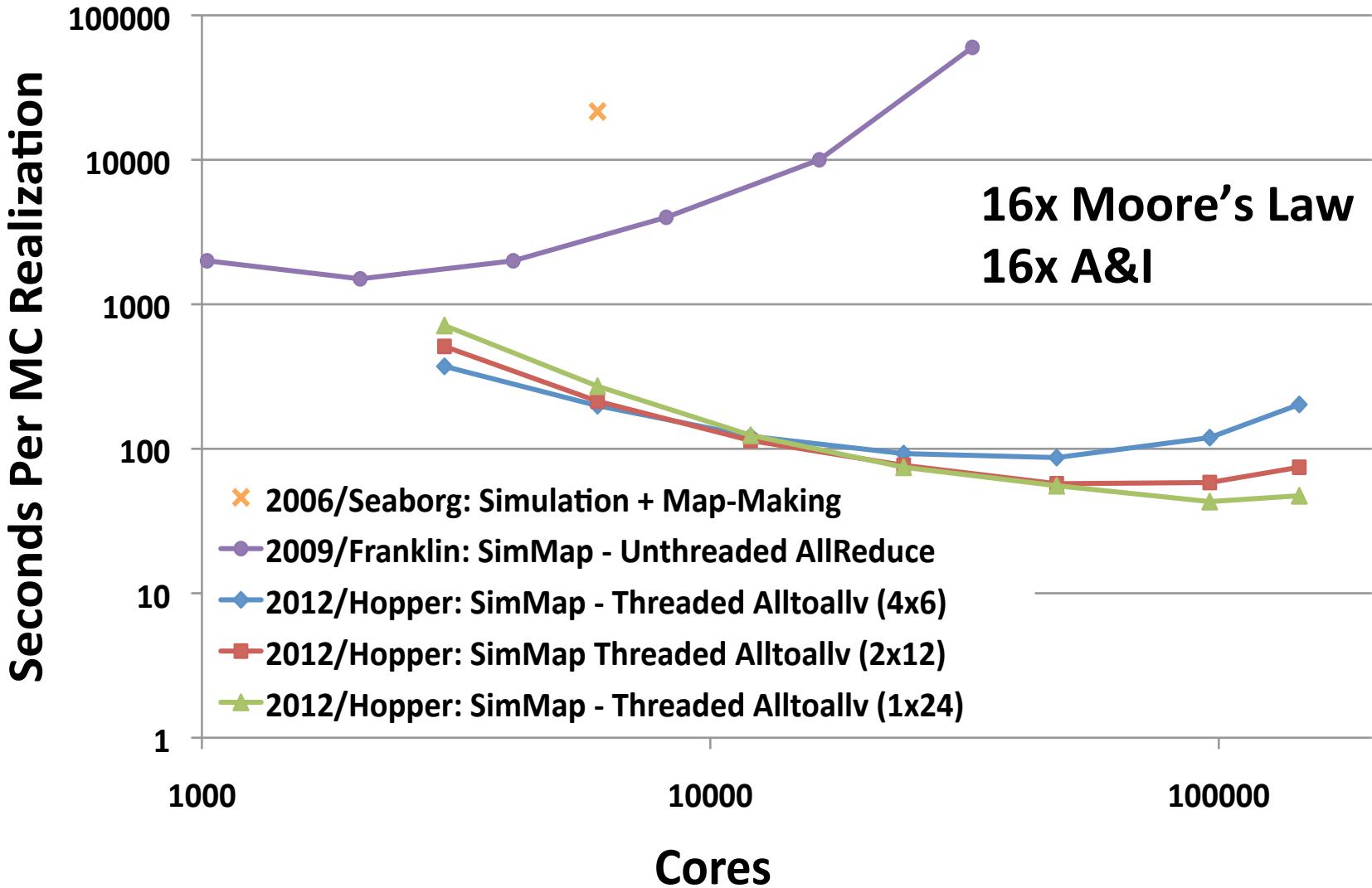
EXACT

- Maximize Gaussian likelihoods
- Scales as $O(\mathcal{N}_p^3)$
- High computational efficiency
- \mathcal{N}_p driven up by
 - high resolution
 - large sky coverage
 - many frequencies
 - polarization
- Intractable for
 - $\mathcal{N}_p > 10^6$ in 2000
 - $\mathcal{N}_p > 10^7$ in 2015
 - $\mathcal{N}_p > 10^8$ in 2030

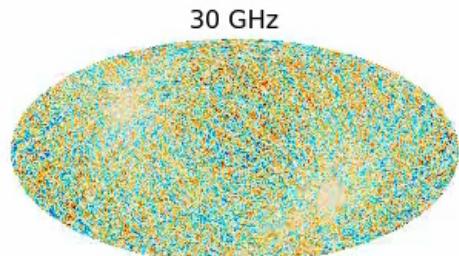
APPROXIMATE

- Monte Carlo replaces covariance
- Scales as $O(\mathcal{N}_t)$
- Low computational efficiency
- \mathcal{N}_t driven up by
 - quest for S/N
 - 10^4 MC realizations
 - many maps per realization
- Intractable for
 - $\mathcal{N}_t > 10^{12}$ in 2000
 - $\mathcal{N}_t > 10^{15}$ in 2015
 - $\mathcal{N}_t > 10^{18}$ in 2030

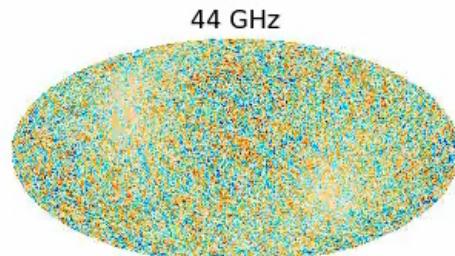
Planck Monte Carlo



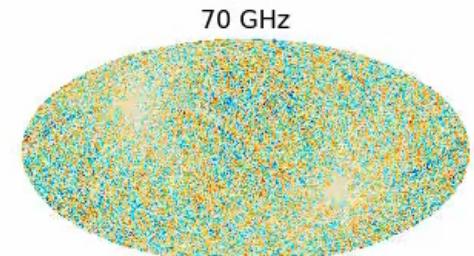
10^4 MC: Planck Noise



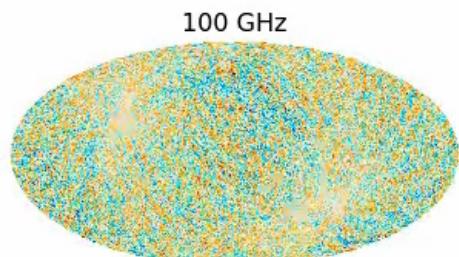
-10 μK 10



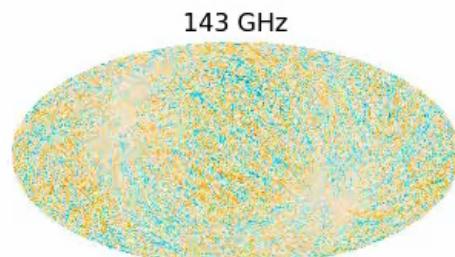
-10 μK 10



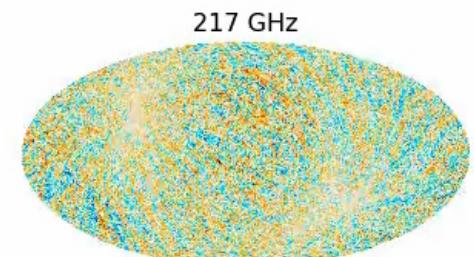
-10 μK 10



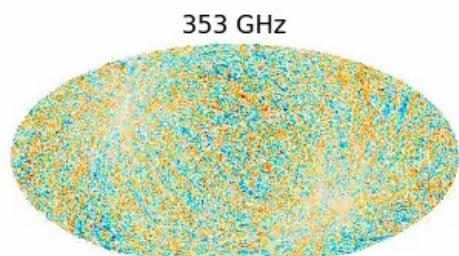
-3 μK 3



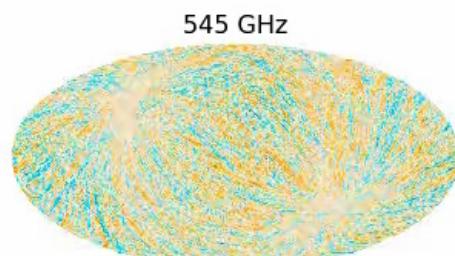
-3 μK 3



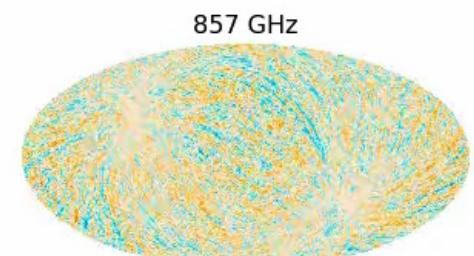
-3 μK 3



-10 μK 10



-0.01 MJy/sr 0.01



-0.01 MJy/sr 0.01

RE 0

~50 million CPU-hours!

Future Challenges

- Precise analysis of forthcoming B-mode experiments requires our algorithms and implementations to track Moore's Law for
 - 10+ years.
 - 3+ generations of supercomputer.
- The means of achieving Moore's Law also evolves
 - Clock speed, node count, accelerators/many-core, ???
- HPC is increasingly energy-constrained
 - Energy-aware algorithms, dark silicon
- Overall: increasing heterogeneity, complexity, imbalance.

Conclusion

Realizing the enormous scientific promise of the coming generations of CMB B-mode experiments will critically depend on our ability to exploit the capabilities of the coming generations of supercomputer, whatever challenges their architectures and energy-constraints may pose.