

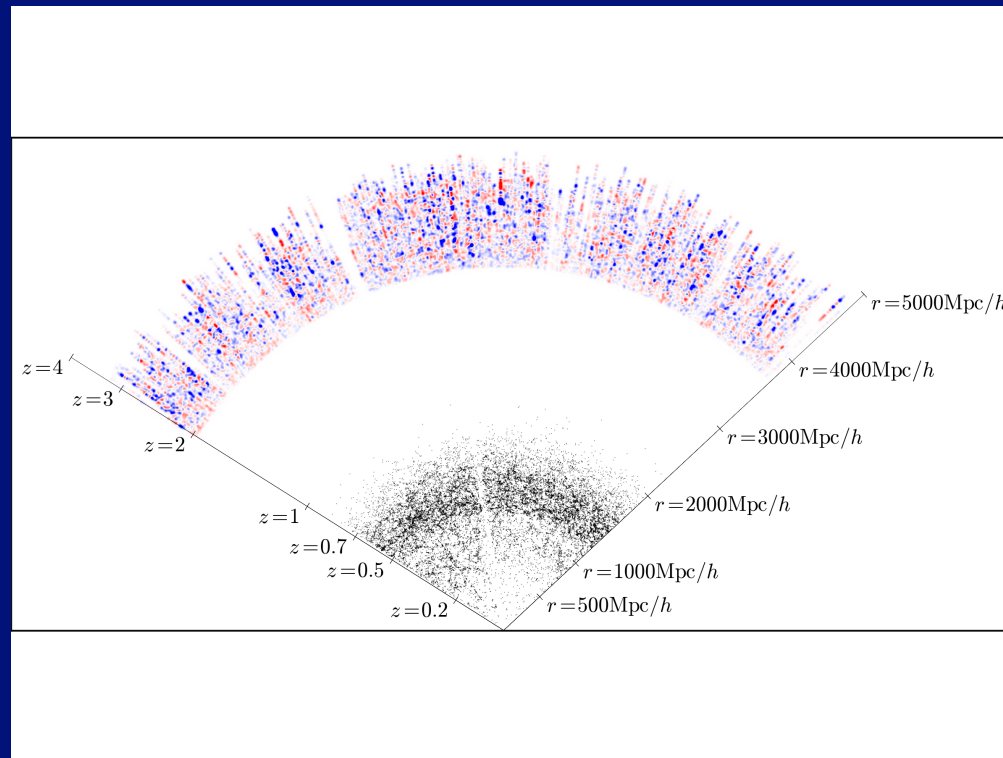
Diffuse Lyman Alpha Emission
at High Redshift:
a Detection using the Cross-Correlation with
QSOs.

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Santander, 18 June 2015

BOSS: spectroscopy of galaxies and quasars

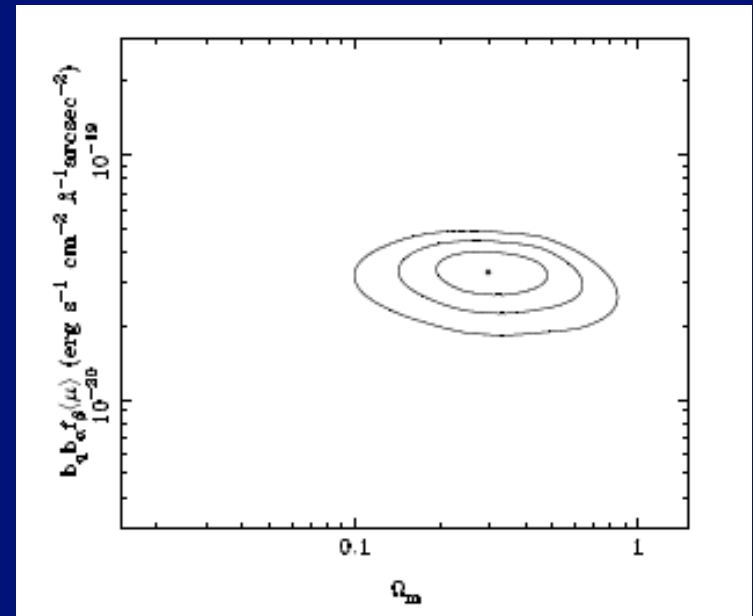
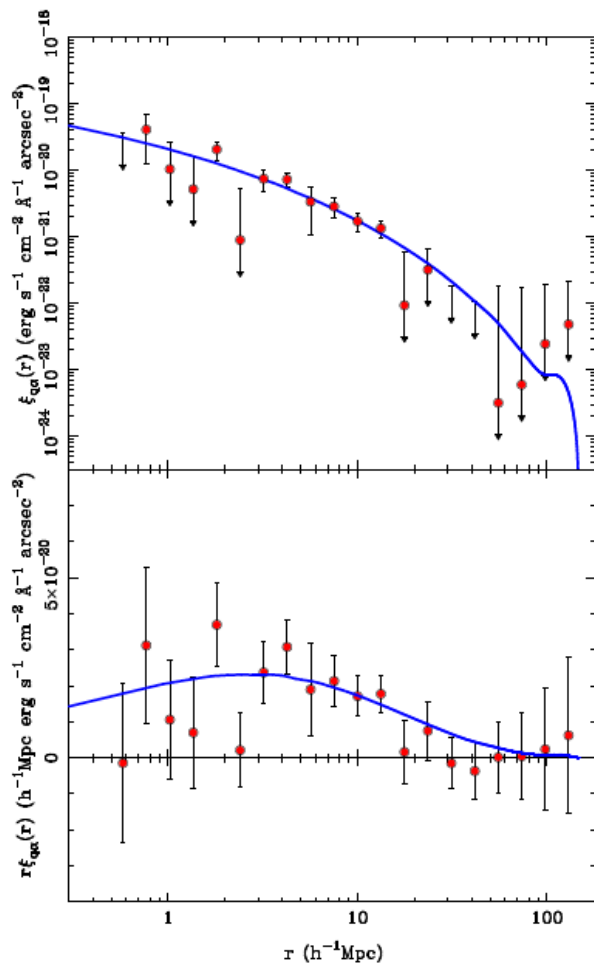


- Most fibers were targeted to massive galaxies at $z < 0.7$, but we can subtract the model galaxy spectrum and use the remaining flux as a tracer of any emission along the line of sight, among that, high- z Ly α emission.

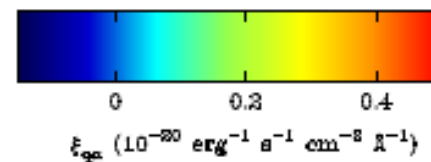
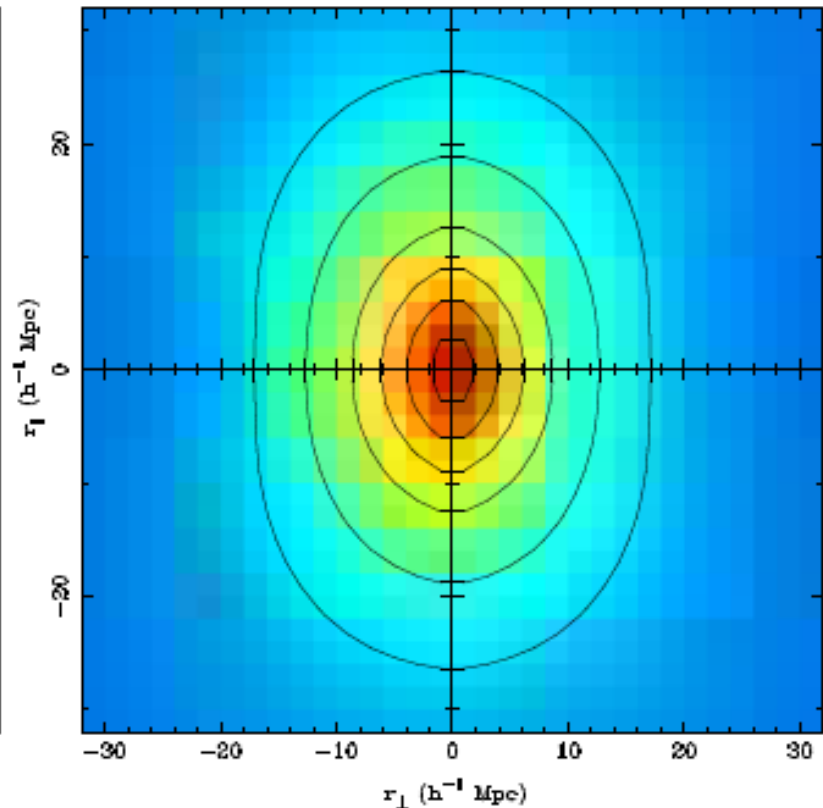
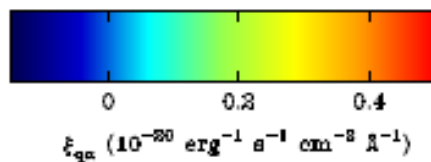
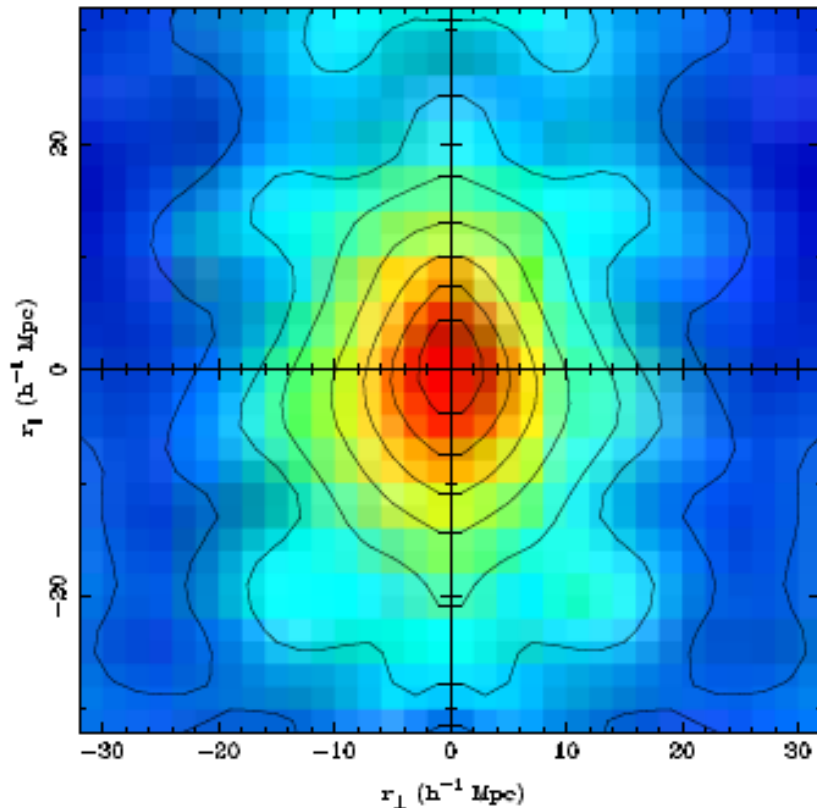
Cross-correlation of Ly α emission with quasars (Croft et al. 2015).

$$b_q b_\alpha f_\beta \langle \mu \rangle = 3.33^{+0.41}_{-0.43} \times 10^{-20} \text{ erg s}^{-1} \text{ cm}^{-2} \text{ \AA}^{-1} \text{ arcsec}^{-2}$$

- The cross-correlation is proportional to the mean Ly α emission, the quasar bias and the bias of Ly α emitters.

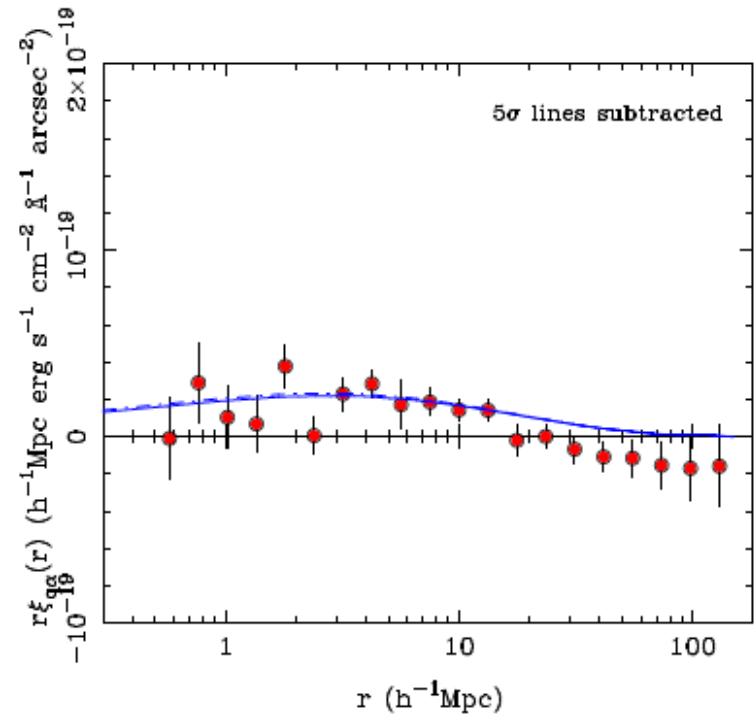
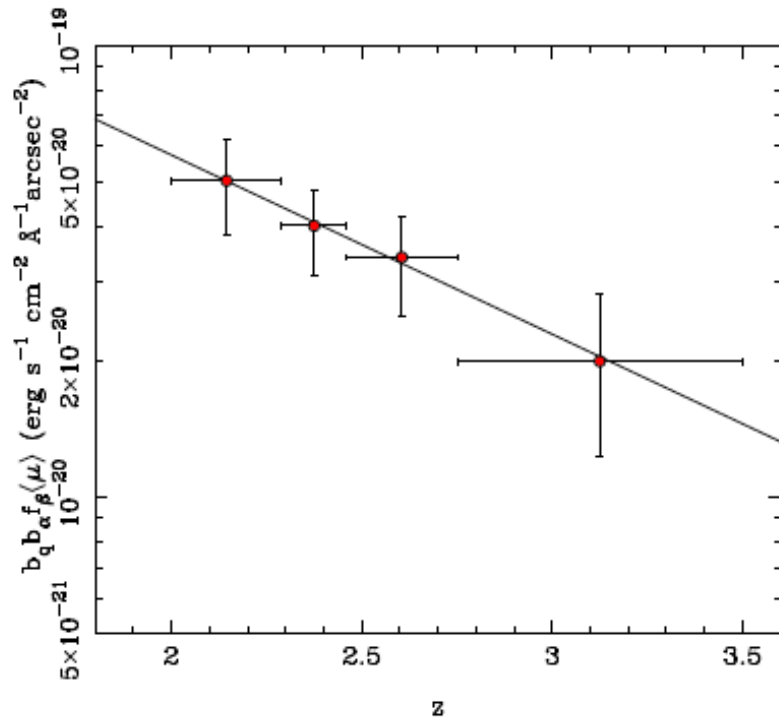


The cross-correlation in redshift space: a radial elongation may originate in radiative transport.



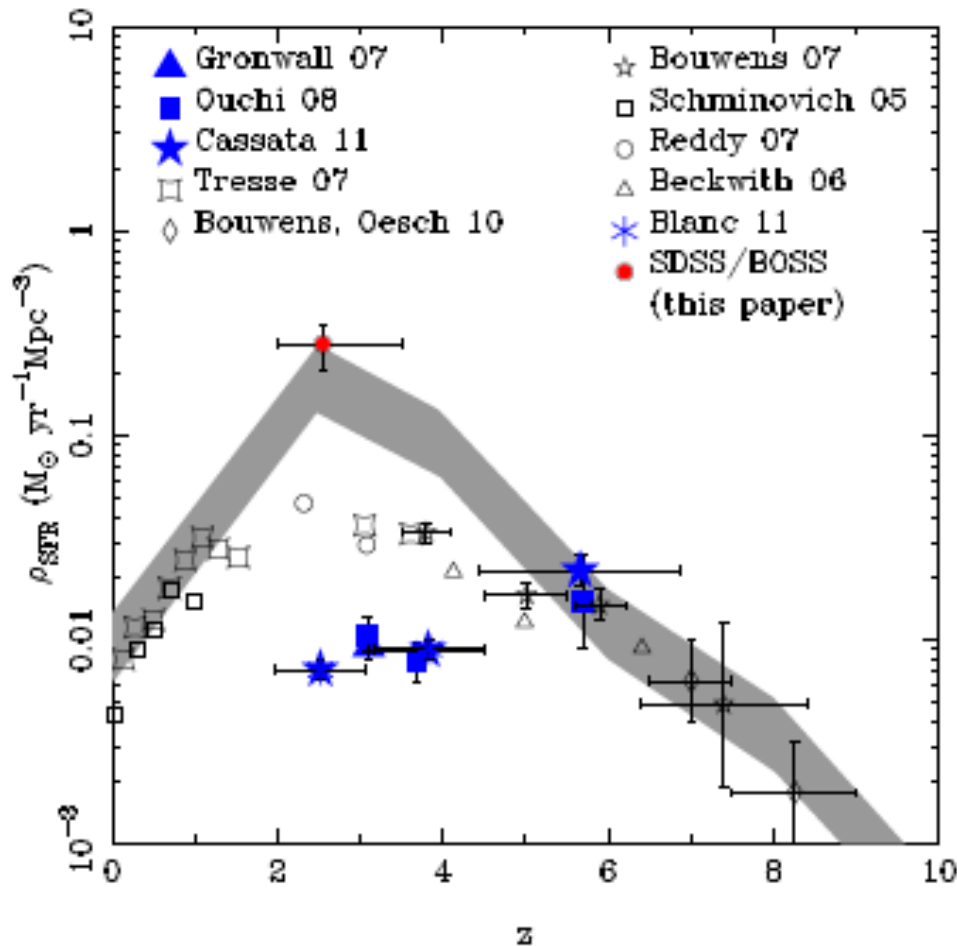
Redshift evolution

No contribution from any highly luminous Ly α galaxies.



A surprisingly high implied star formation rate density:

$$b_q b_\alpha f_\beta \mu_\alpha = (3.3 \pm 0.4) \times 10^{-20} \text{ erg/s/cm}^2 / \text{A} / \text{arcsec}^2$$



Implication: for $b_\alpha=3$ and the observed quasar bias $b_q=3.6$, the inferred Ly α emission is nearly all the observed star formation rate density after correcting for dust absorption!

This is surprising: it implies that even though most UV continuum photons are absorbed by dust, most Ly α photons escape from large gaseous halos of low Ly α surface brightness surrounding galaxies.

The only other possible source of this Ly α emission is cooling radiation from gas in halos, but still short by a factor ~ 10 to account for this measurement.

Conclusions

- We have measured the cross-correlation of Ly α brightness and quasars. The result is an 8- σ detection at $z=2.5$, interpreted as the clustering of star-forming galaxies around quasars. The surprising implication is that most Ly α escape out to gaseous halos of low surface brightness around galaxies, from where they are emitted into the intergalactic medium.